

Recall the differential equations of planetary motion

$$\ddot{r} - r \dot{\theta}^2 = -\frac{\mu}{r^2} \quad (1) \quad \text{where } \mu = GM$$

$$2 \dot{r} \dot{\theta} + r \ddot{\theta} = 0 \quad (2)$$

Consider the expression $\dot{r}(1) + r \dot{\theta}(2)$:

$$\dot{r} \ddot{r} - r \dot{r} \dot{\theta}^2 + 2 r \dot{r} \dot{\theta}^2 + r^2 \dot{\theta} \ddot{\theta} = -\frac{\mu}{r^2} \dot{r}$$

This is equivalent to

$$\frac{d}{dt} \left(\frac{1}{2} \dot{r}^2 \right) + \frac{d}{dt} \left(\frac{1}{2} r^2 \right) \dot{\theta}^2 + \cancel{r^2 \frac{d}{dt} \left(\frac{1}{2} \dot{\theta}^2 \right)} = \mu \frac{d}{dt} \left(\frac{1}{r} \right)$$

Combining all the terms together we have

$$\begin{aligned} \frac{d}{dt} \left(\frac{1}{2} \dot{r}^2 - \mu \frac{1}{r} \right) + \frac{d}{dt} \left(\frac{1}{2} r^2 \dot{\theta}^2 \right) \\ = \frac{d}{dt} \left(\frac{1}{2} \dot{r}^2 + \frac{1}{2} r^2 \dot{\theta}^2 - \frac{\mu}{r} \right) = 0 \end{aligned}$$

$$\text{So } E = \underbrace{\frac{1}{2} \dot{r}^2 + \frac{1}{2} r^2 \dot{\theta}^2}_{\text{Kinetic Energy}} - \underbrace{\frac{\mu}{r}}_{\text{Potential Energy}} \text{ is a constant}$$

Recall that the velocity $\frac{d\vec{r}}{dt} \equiv \vec{v} = \dot{r} \hat{r} + r \dot{\theta} \hat{\theta}$

$$\begin{aligned} \text{so } \|\vec{v}\|^2 &= (\dot{r} \hat{r} + r \dot{\theta} \hat{\theta})^2 = \dot{r}^2 \hat{r} \cdot \hat{r} + 2 r \dot{r} \dot{\theta} \hat{r} \cdot \hat{\theta} + r^2 \dot{\theta}^2 \hat{\theta} \cdot \hat{\theta} \\ &= \dot{r}^2 + r^2 \dot{\theta}^2 \end{aligned}$$

Hence $K.E = \frac{1}{2} \|\vec{v}\|^2$ as in many other systems