

# Complex trajectories of a simple pendulum

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# Work based on:

- Carl M. Bender, Darryl D. Holm, Daniel W. Hook (2007) “Complex Trajectories of a Simple Pendulum”, Journal of Physics A40, F81-F89, [[math-ph/0609068](#)]
- Carl M. Bender, Darryl D. Holm, Daniel W. Hook (2007) “Complexified Dynamical Systems”, to appear: Journal of Physics A, [[arXiv:0705.3893](#)]
- Carl M. Bender, Jun-Hua Chen, Daniel W. Darg, Kimball A. Milton “Classical Trajectories for Complex Hamiltonians” (2006) Journal of Physics A39, 4219-4238 [[math-ph/0602040](#)]
- Carl M. Bender, Daniel W. Darg (2007) “Spontaneous Breaking of Classical PT Symmetry”, [[hep-th/0703072](#)]

# The Simple Pendulum

Position:

$$X = L \sin \theta$$

$$Y = -L \cos \theta$$

Velocity:

$$\dot{X} = L \dot{\theta} \cos \theta$$

$$\dot{Y} = L \dot{\theta} \sin \theta$$

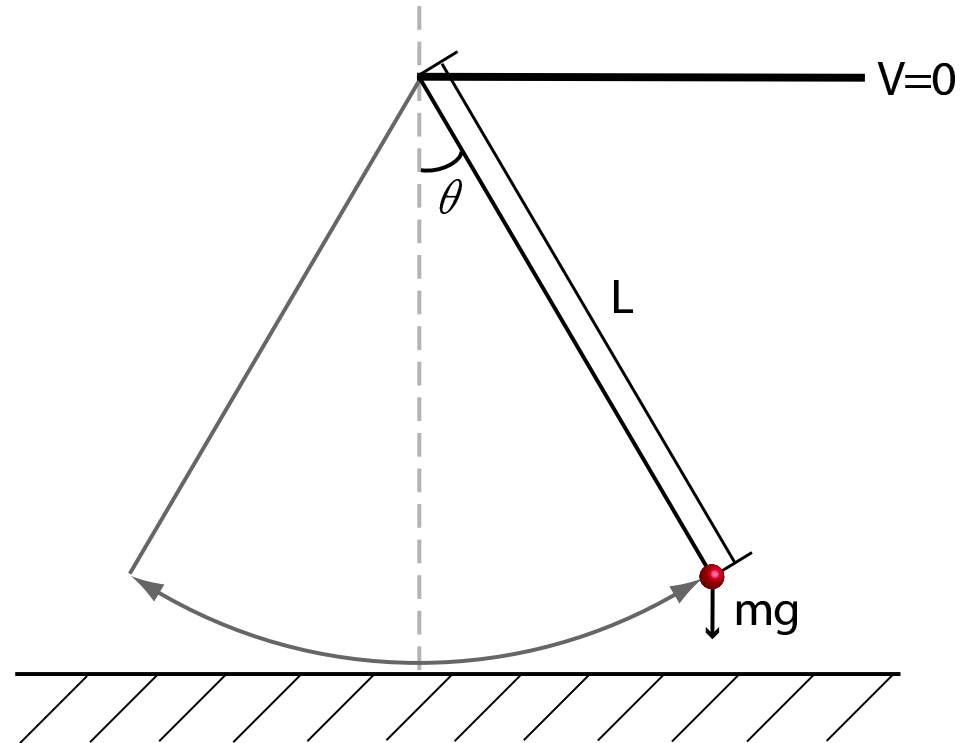
Energies:

$$V = -mgL \cos \theta$$

$$T = \frac{1}{2}m(\dot{X}^2 + \dot{Y}^2) = \frac{1}{2}mL^2\dot{\theta}^2$$

Hamiltonian:

$$H = \frac{1}{2}mL^2\dot{\theta}^2 - mgL \cos \theta$$



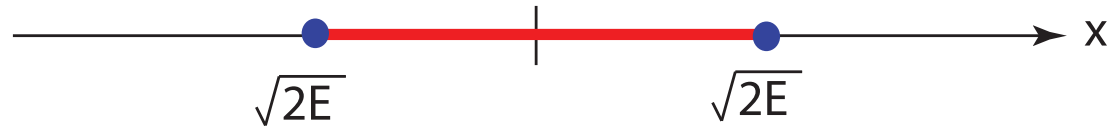
# Turning points

We have turning points in the motion when

$$p = 0$$

Hence,

$$H = V(x) = E$$



For a simple harmonic oscillator with Hamiltonian

$$H = \frac{1}{2}p^2 + x^2$$

$$x_0 = \sqrt{2E}$$

# Complexifying the Simple pendulum

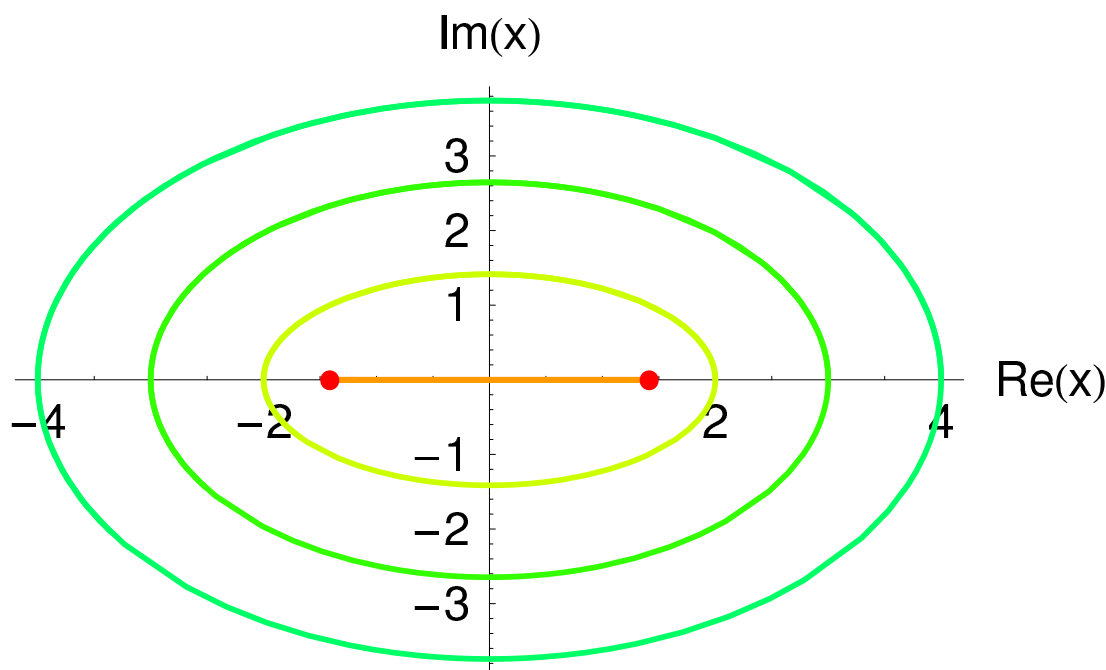
Gravity:

$$g = 1$$

Turning points:

$$x_0 = \sqrt{2E}$$

There is a branch cut  
between turning points  
( $x = \sqrt{2E}$ )



Paths have the same  
period:

$$T = \oint_C dx / \sqrt{2[E - V(x)]}$$

# A periodic potential

Case 1:  $-1 \leq E \leq 1$  (Swinging pendulum)

Hamiltonian:

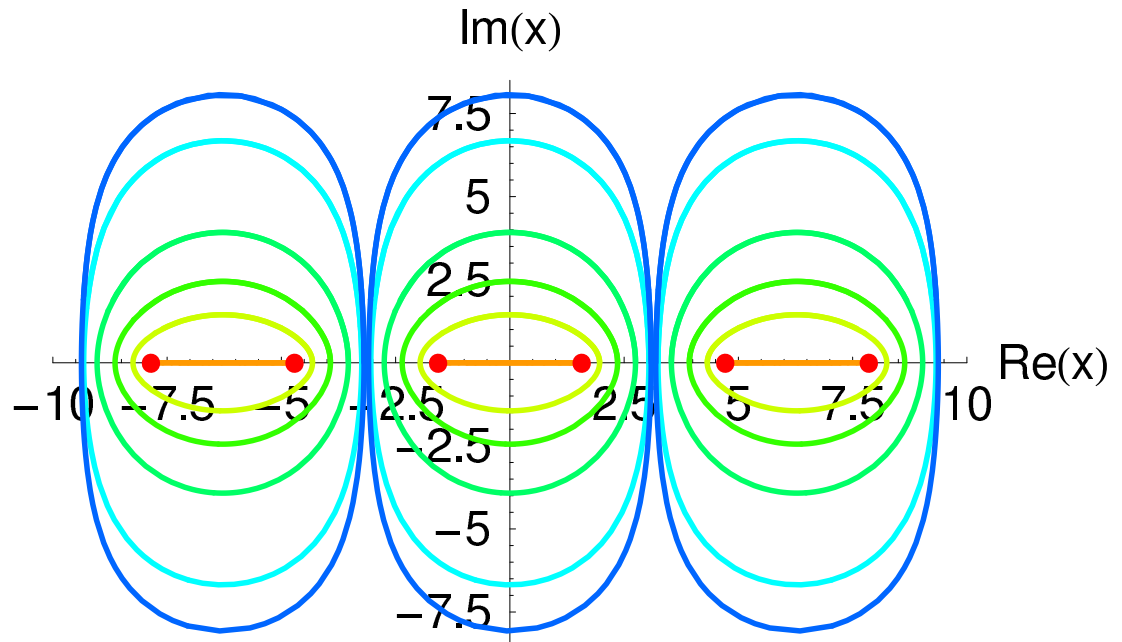
$$H = \frac{1}{2}p^2 - \cos x$$

Gravity:

$$g = 1$$

Turning points:

$$x_0 = \pi/2 + n\pi$$



# Example: Real Gravitational Field

Case 2:  $|E| \geq 1$  (Rotating pendulum)

Hamiltonian:

$$H = \frac{1}{2}p^2 - \cos x$$

Gravity:

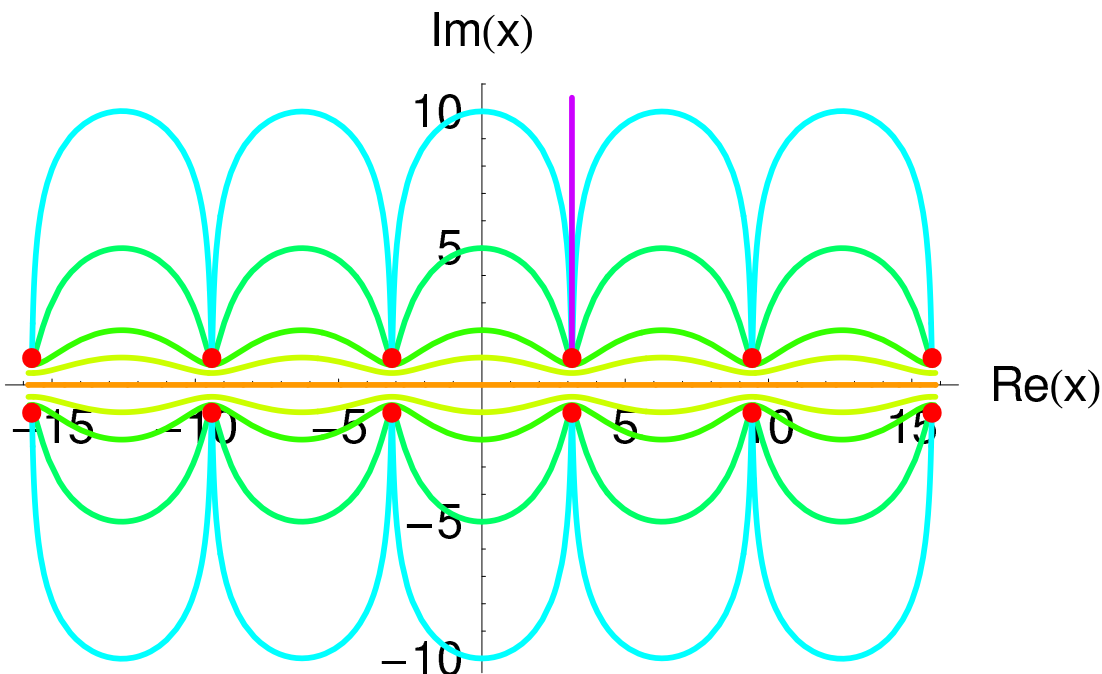
$$g = 1$$

Energy:

$$E = \cosh 1$$

Turning points:

$$x_0 = (2k + 1)\pi \pm i$$

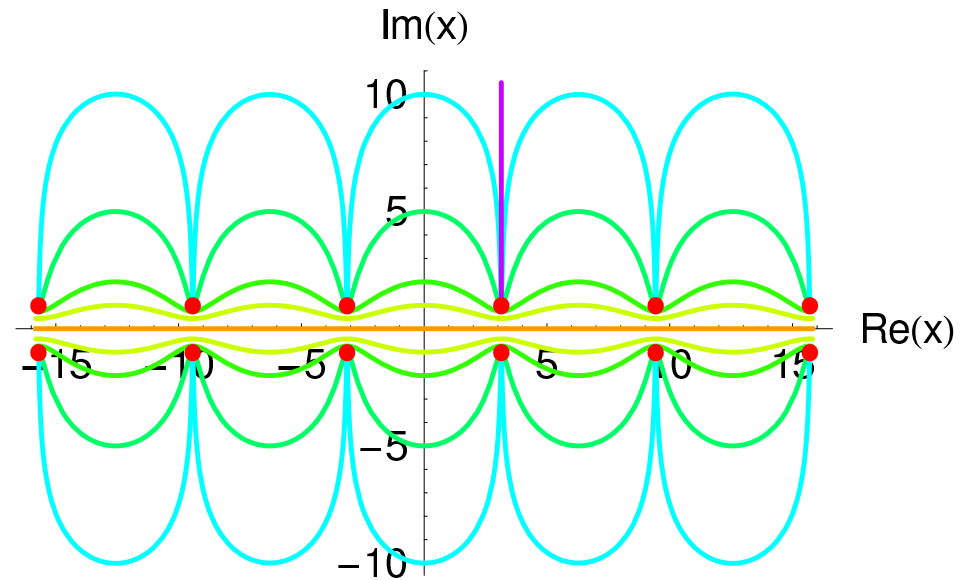


# Example: Real Gravitational Field

Time to follow the purple path . . .

$$T = \frac{1}{\sqrt{2}} \int_{x=i+\pi}^{i\infty+\pi} \frac{dx}{\sqrt{E + \cos x}}$$

$$T = 1.97536 \dots$$



# Example: Real Gravitational Field

Case 3:  $E \leq -1$  (Unphysical pendulum)

Hamiltonian:

$$H = \frac{1}{2}p^2 - \cos x$$

Gravity:

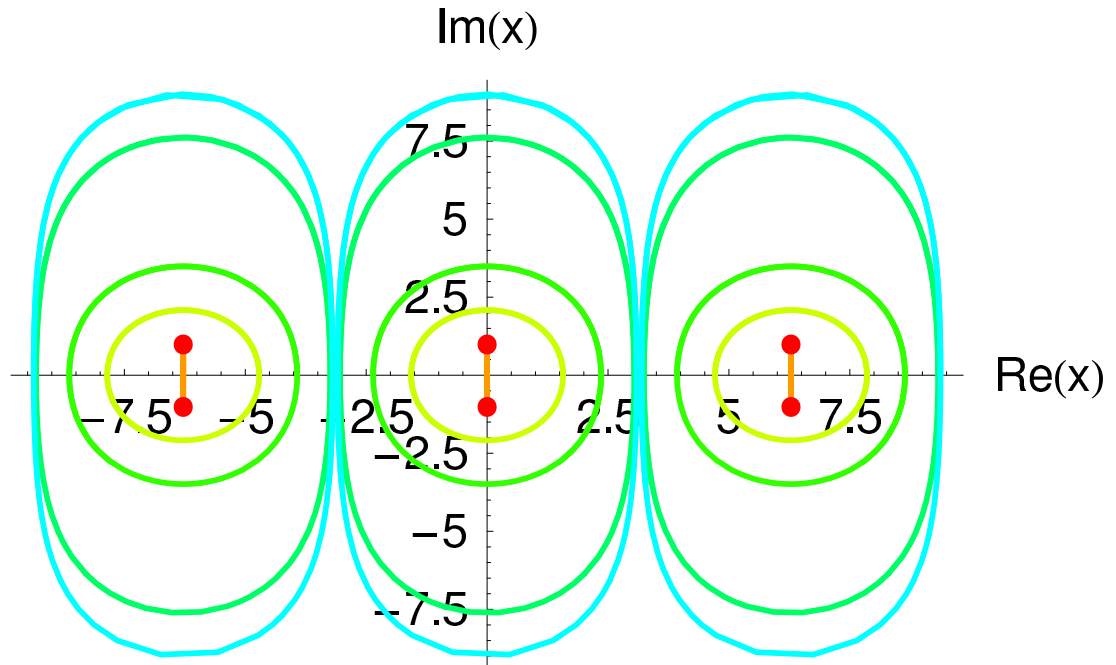
$$g = 1$$

Energy:

$$E = -\cosh 1$$

Turning points:

$$x_0 = 2\pi k + \pm i$$



# Example: Imaginary Gravitational Field

Case 1:  $E = \sinh 1$

Hamiltonian:

$$H = \frac{1}{2}p^2 - i \cos x$$

Gravity:

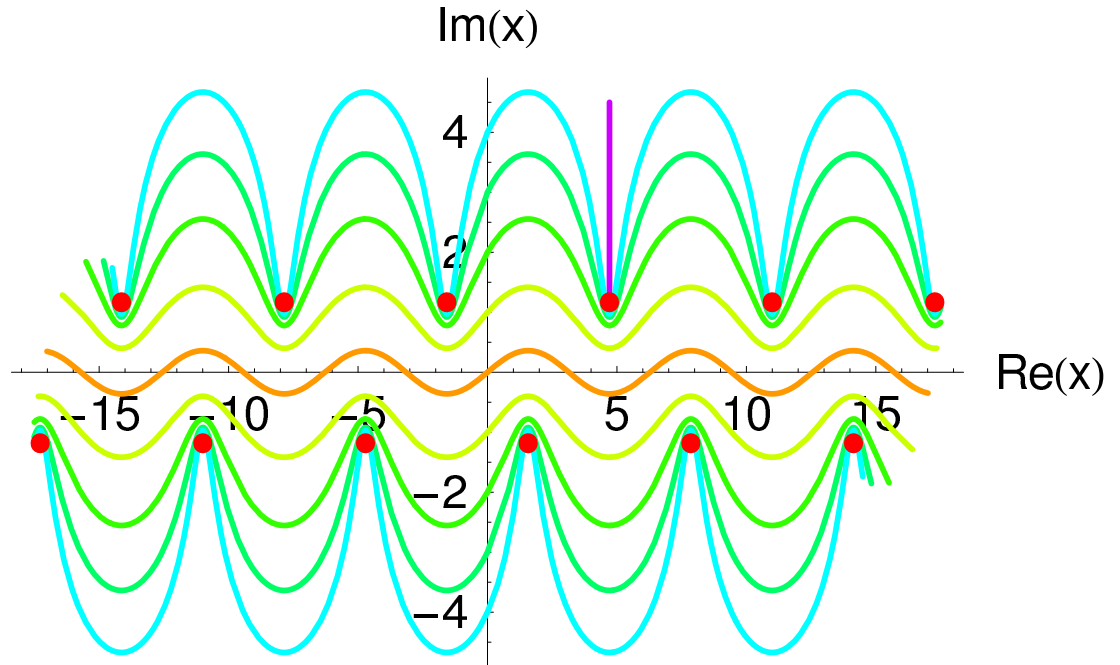
$$\mathbf{g} = \mathbf{i}$$

Energy:

$$\mathbf{E} = \sinh 1$$

Turning points:

$$x_0 = \left(n + \frac{1}{2}\right)\pi + (-1)^n i$$

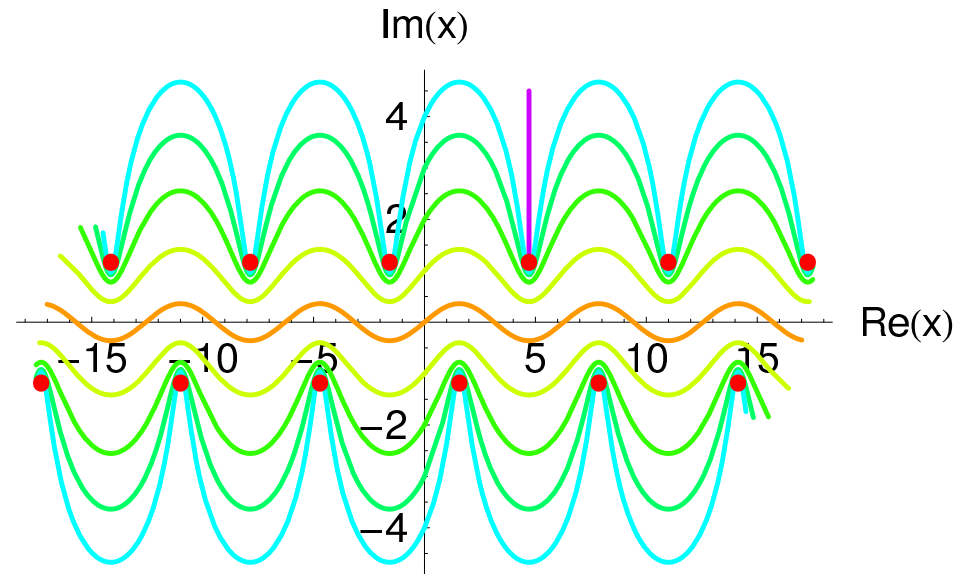


# Example: Imaginary Gravitational Field

Time to follow the purple path . . .

$$T = \frac{1}{\sqrt{2}} \int_{x=i+3\pi/2}^{i\infty+3\pi/2} \frac{dx}{\sqrt{E + i \cos x}}$$

$$\begin{aligned} T &= \frac{1}{\sqrt{2}} \int_{s=1}^{\infty} \frac{ds}{\sqrt{\sinh s - \sinh 1}} \\ &= \frac{2}{\sqrt{e}} K(-1/e^2) = 1.84549 \dots \end{aligned}$$



# Example: Imaginary Gravitational Field

Case 2:  $E = -\sinh 1$

Hamiltonian:

$$H = \frac{1}{2}p^2 - i \cos x$$

Gravity:

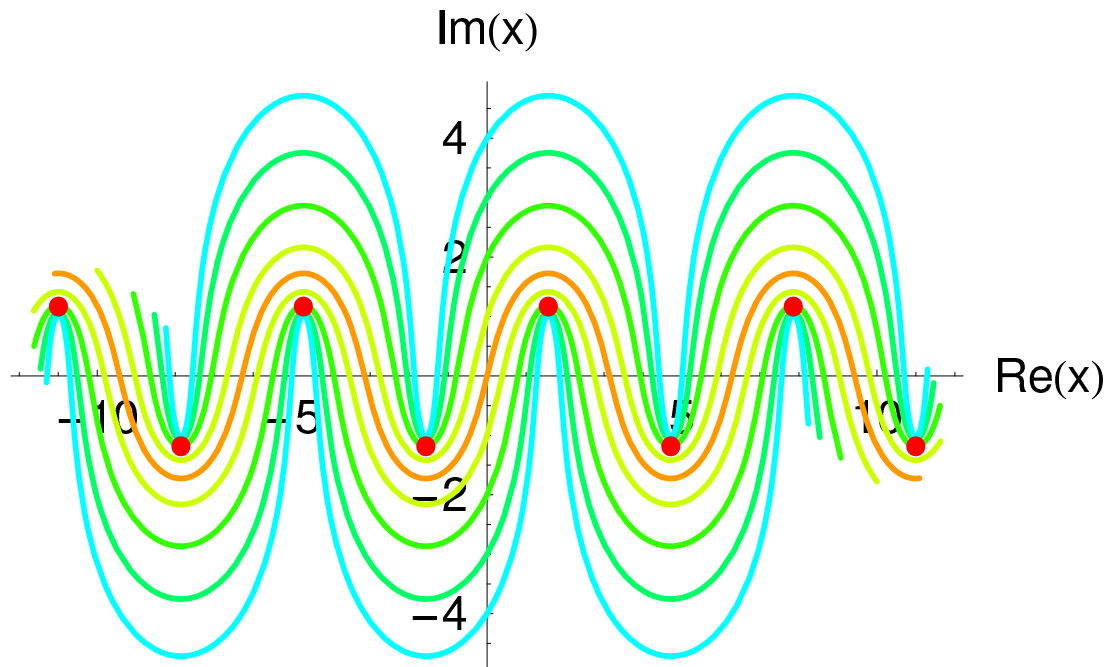
$$\mathbf{g} = \mathbf{i}$$

Energy:

$$\mathbf{E} = -\sinh 1$$

Turning points:

$$x_0 = \left(n + \frac{1}{2}\right)\pi + (-1)^{n+1}i$$

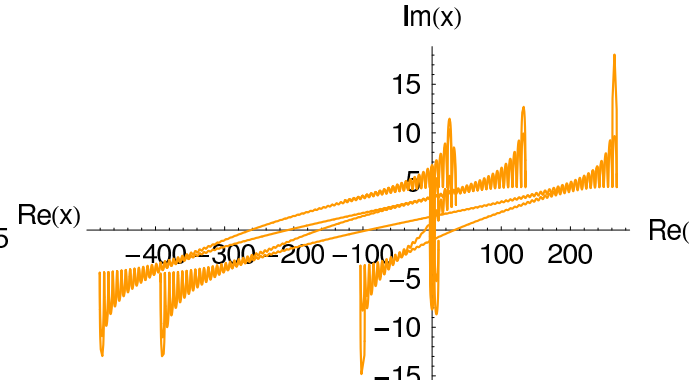
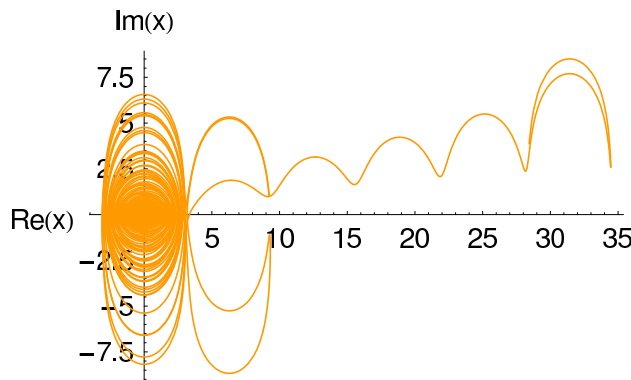
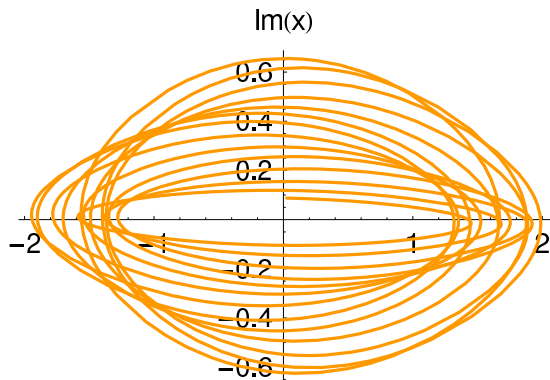


# What about chaotic systems?

Introduce a driving term of the form  $\epsilon \sin(\omega t)$ :

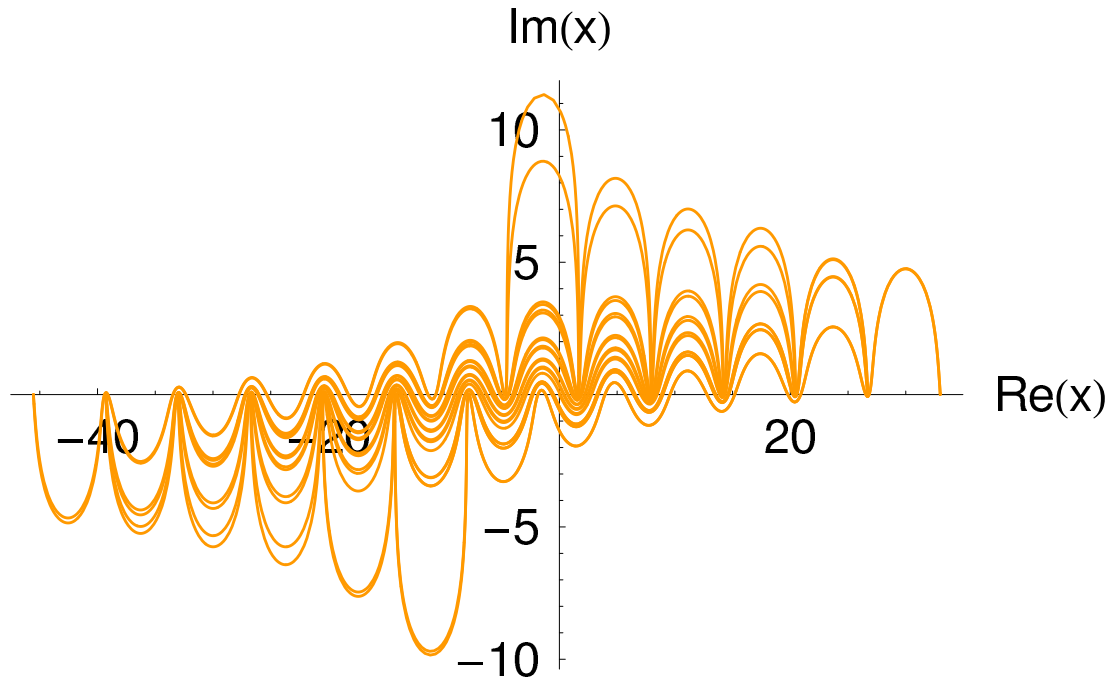
$$H = \frac{1}{2}p^2 - g \cos x - \epsilon \sin(\omega t)$$

$$g = i, \epsilon > 0$$



# Imaginary energy?

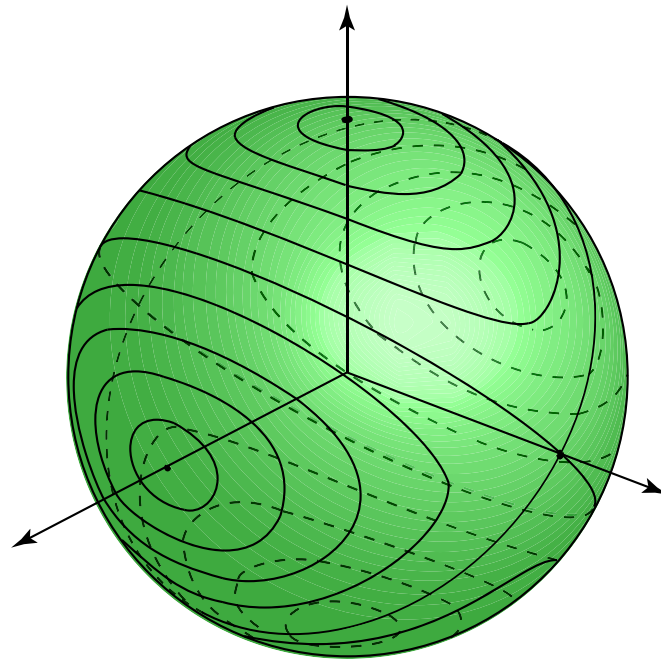
$g = i$  and  $E = i$



# Euler Equations

The Euler equations governing the rotation of a rigid body about a fixed axis also are  $\mathcal{PT}$ .

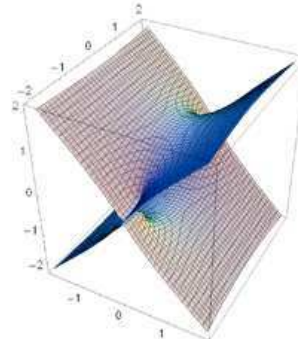
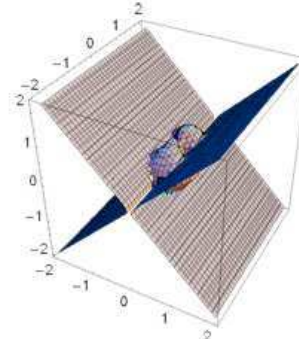
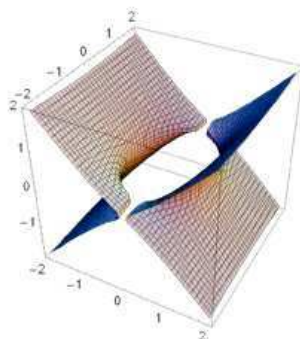
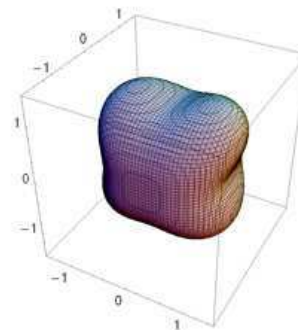
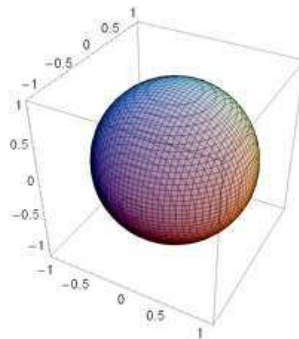
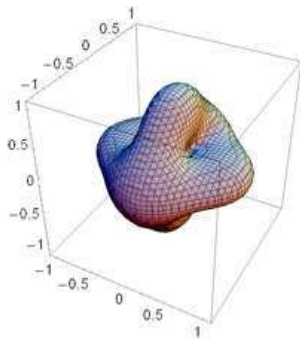
$$\begin{aligned}\dot{L}_1 &= L_2 L_3 \\ \dot{L}_2 &= -2L_1 L_3 \\ \dot{L}_3 &= L_1 L_2\end{aligned}$$



# Euler Equations

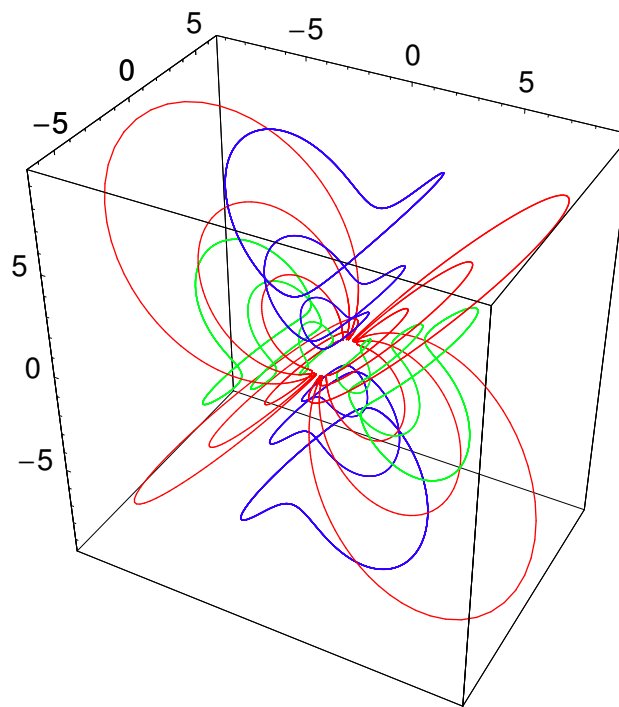
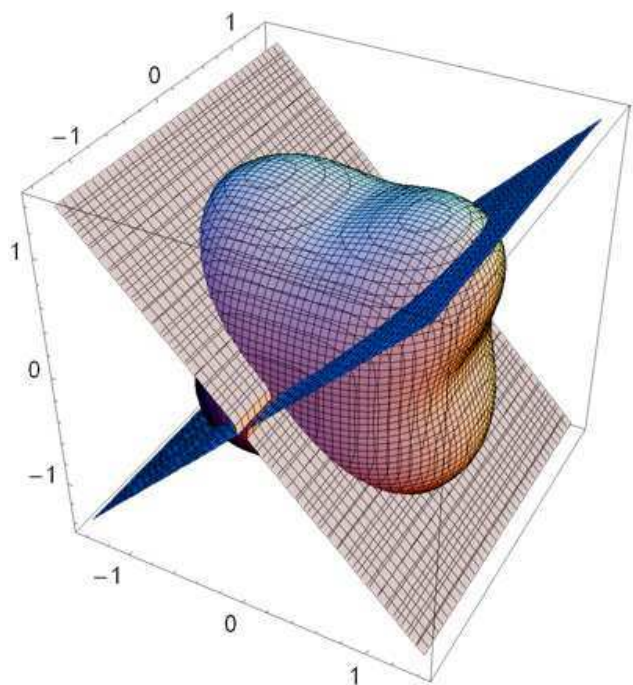
We extend the Euler Equations into the complex plane. There are two constants of motion (which gives us 4 constraints):

$$C = \frac{1}{2} (x_1^2 + x_2^2 + x_3^2) \quad \text{and} \quad H = \frac{1}{2} x_3^2 - \frac{1}{2} x_1^2.$$



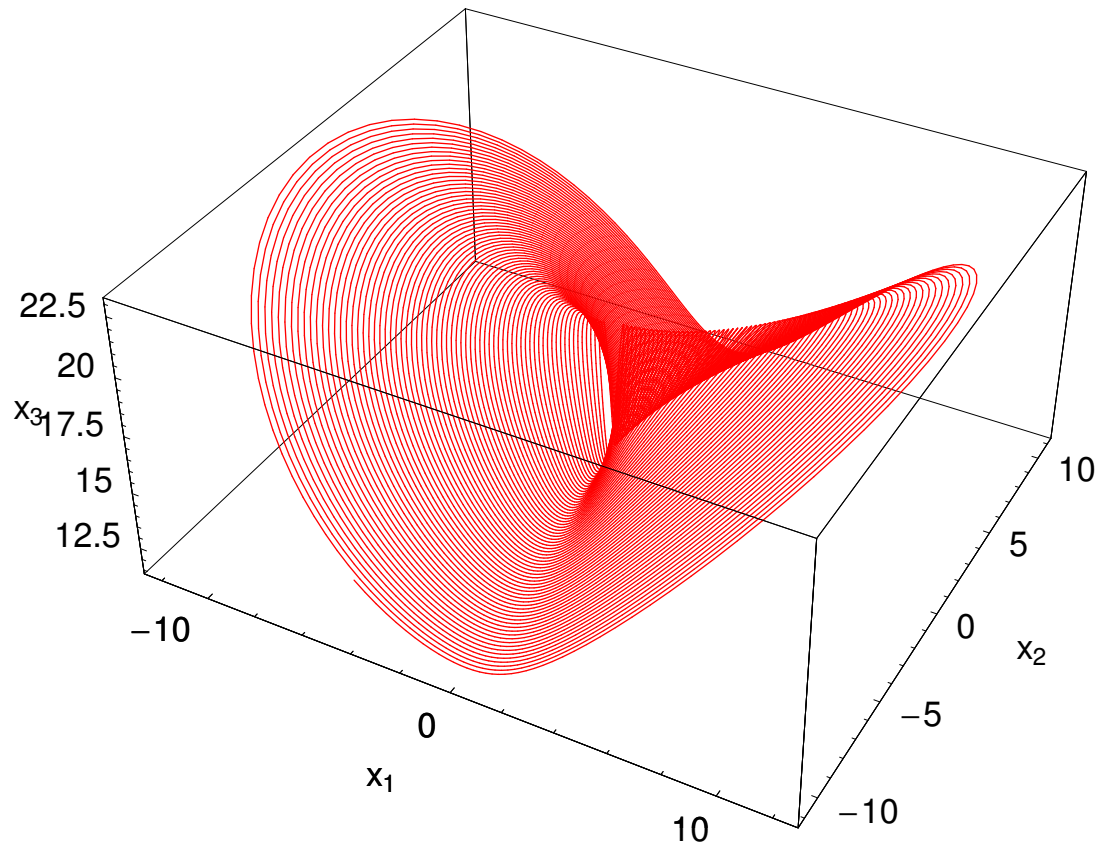
# Euler Equations

Intersections of the level surfaces give us the orbits:



# Euler Equations

If we allow energy to be imaginary:



# Lotka-Volterra equations

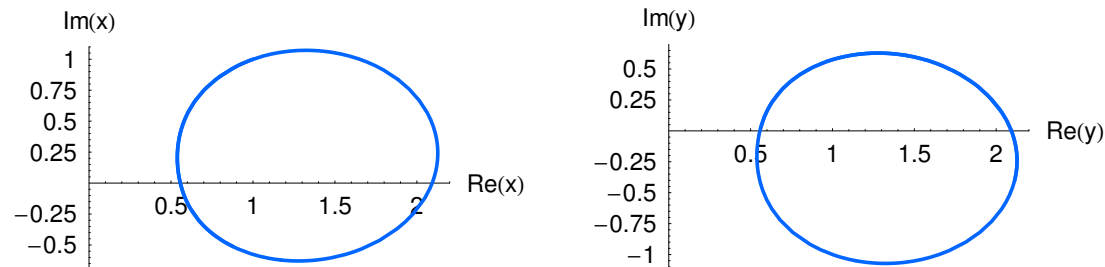
Another PT-Symmetric system - this time under:

$$\dot{x} = x - xy, \quad \dot{y} = -y + xy \quad (1)$$

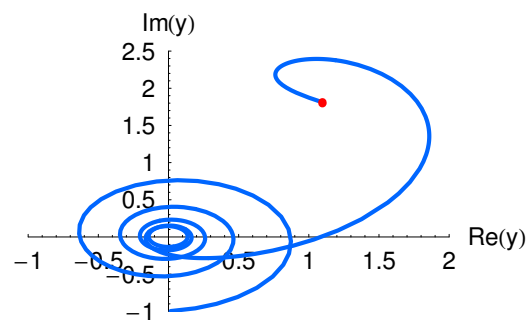
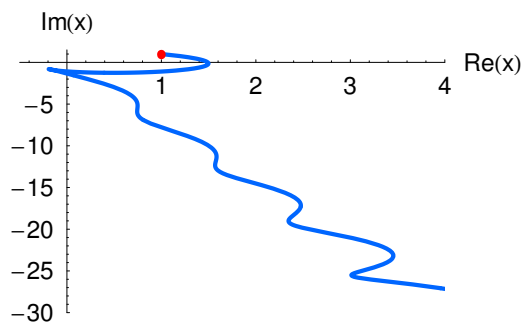
constant of motion

$$x + y - \log(xy) = C. \quad (2)$$

Fox/ rabbit interchange, and time reversal.



# Lotka-Volterra equations



# Conclusions

- Classical Mechanics seems to naturally extend into the complex plane to give some familiar and not so familiar results
- Complex extensions of classical mechanics help us understand well known systems

# Where next?

- spherical pendulum, the spinning top
- other discrete dynamical systems - SIR models etc.
- classification of chaotic systems using  $\mathcal{PT}$  Symmetry (see for example, the Kicked Rotor).